Xtreme QCD

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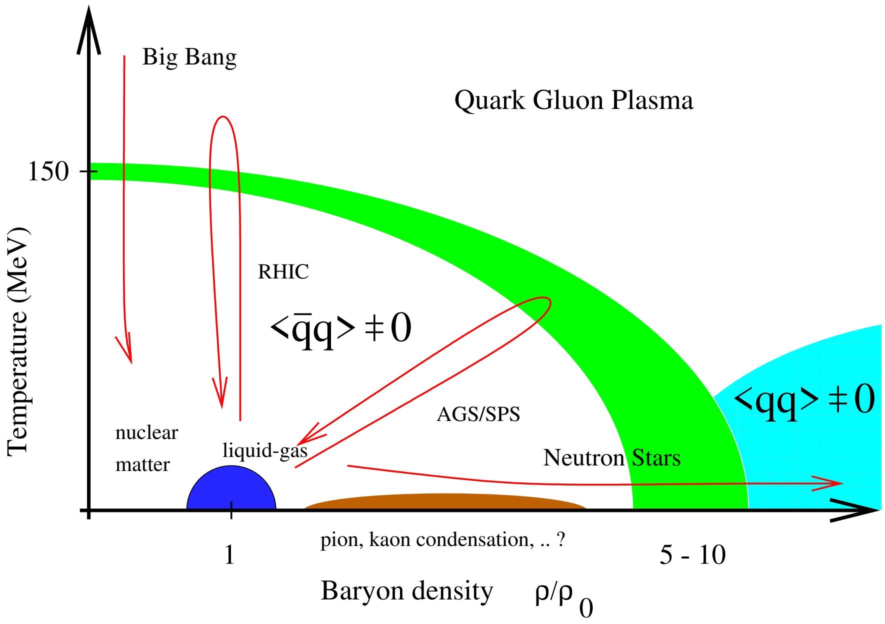
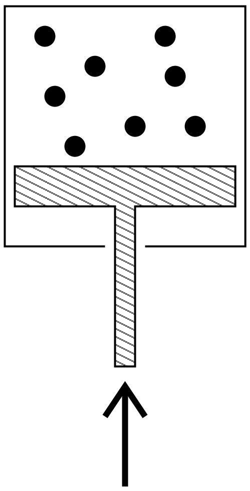
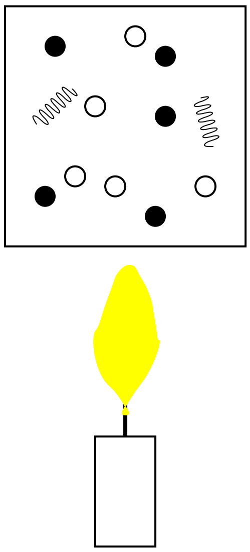


QCD in Xtreme Conditions

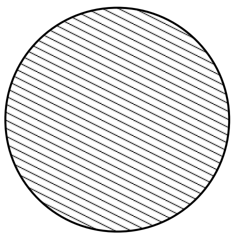
High Density



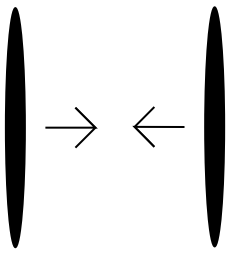
High Temperature



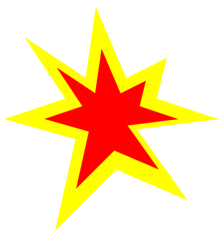
more practical:



neutron
stars



relativistic
heavy–ion
collisions



big bang

WHY?

- These conditions exist in the universe:

  Neutron Stars, Big Bang

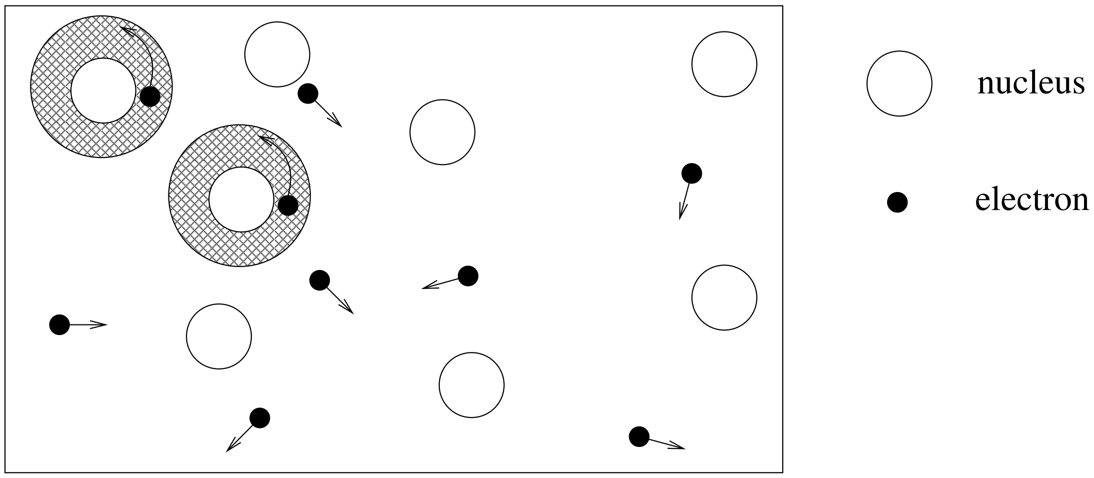
- Exploring the entire phase diagram is important to understanding the phase that we happen to live in

  We cannot understand the structure of hadrons without understanding the non-trivial QCD vacuum.
  We cannot understand the vacuum state without understanding how it can be modified.

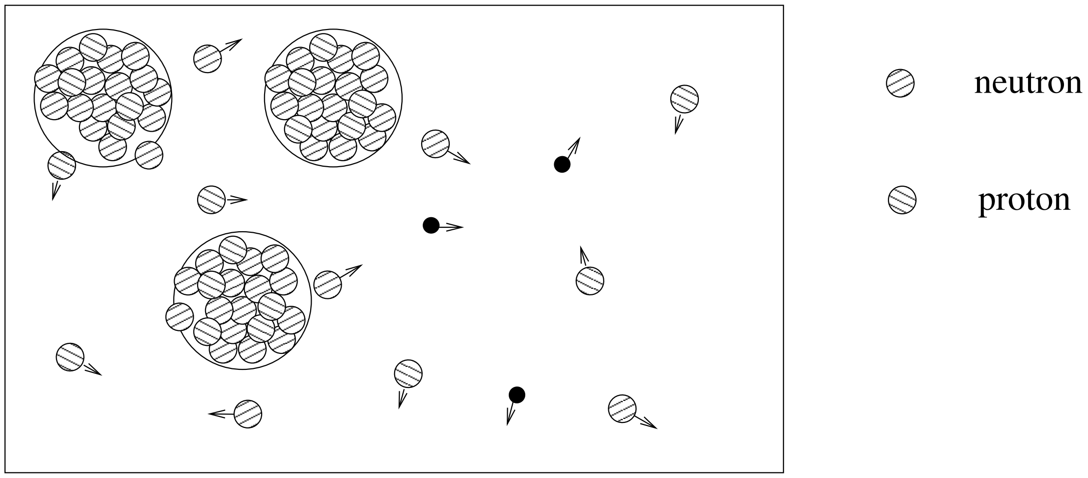
- QCD simplifies in extreme environments:

  Study QCD matter in a regime where quarks and gluons are indeed the correct degrees of freedom.

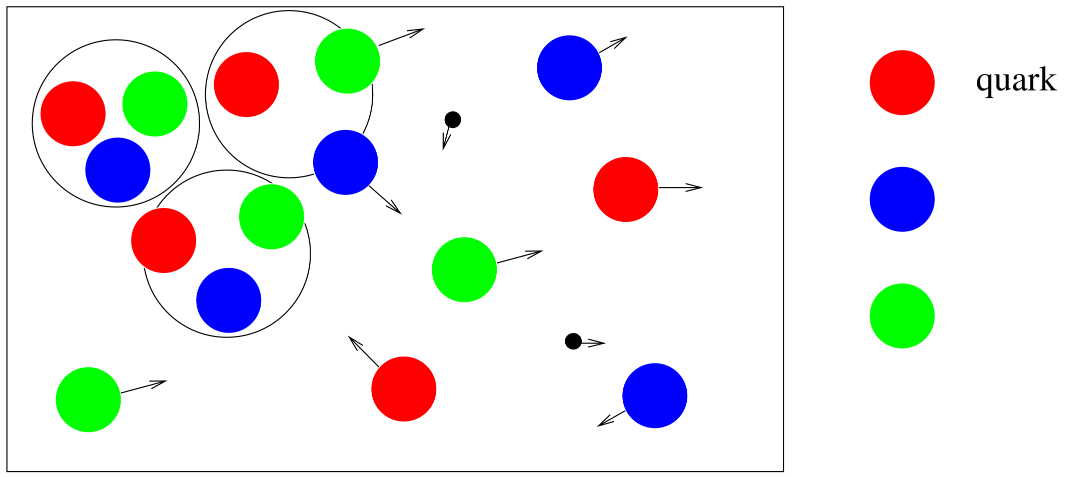
Dense Matter



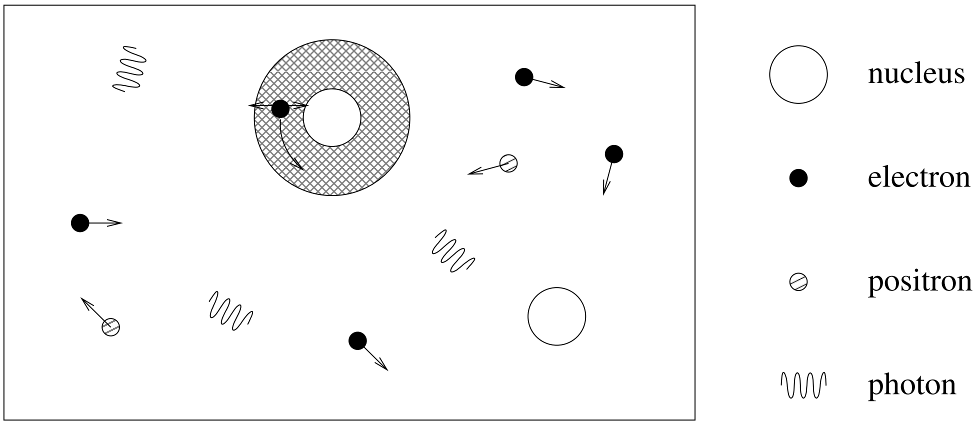
white dwarf matter



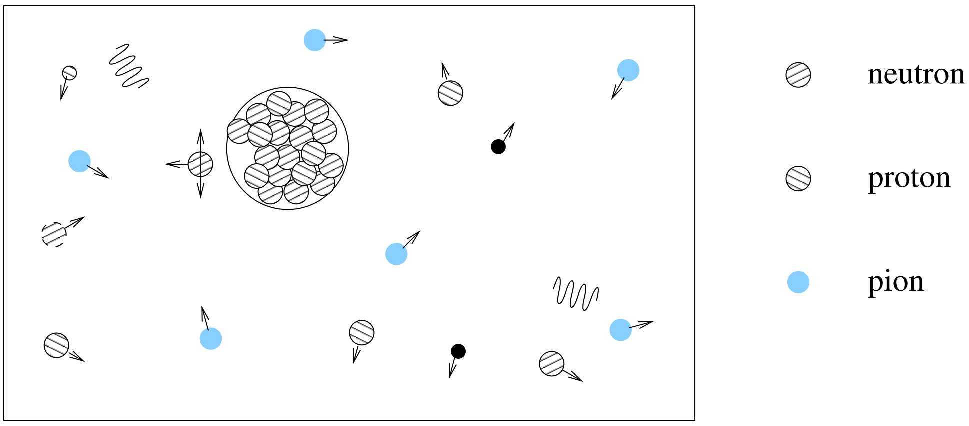
neutron matter



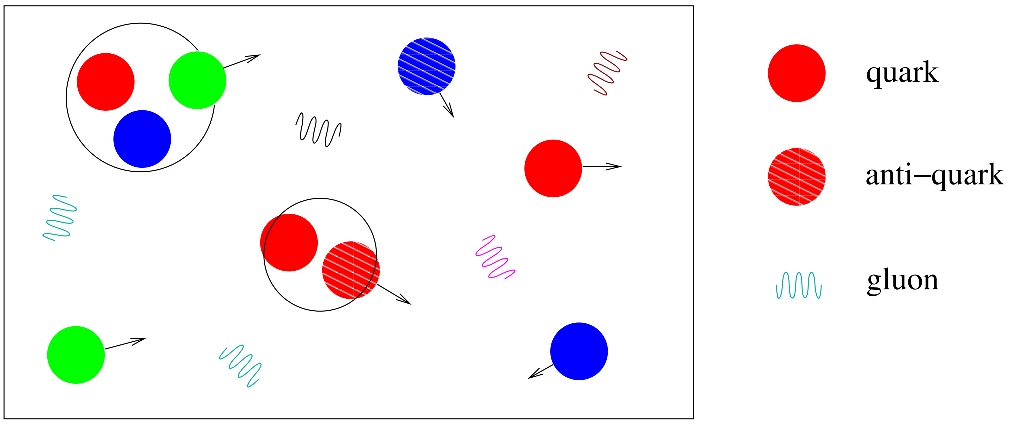
Hot Matter



Plasma



Hadron Gas



Quark Gluons Plasma

Quantum Chromo Dynamics

- elementary fields

$$\text{quarks} \qquad (q_\alpha)^a_f \qquad \begin{cases} \text{color} & a = 1,\ldots,3 \\ \text{spin} & \alpha = 1,2 \\ \text{flavor} & f = u,d,s,c,b,t \end{cases}$$

$$\text{gluons} \qquad A^a_\mu \qquad \begin{cases} \text{color} & a = 1,\ldots,8 \\ \text{polarization} & \epsilon^\pm_\mu \end{cases}$$

- dynamics: gauge theory

$$\mathcal{L} = \sum_f^{N_f} \bar{q}_f (i\not{D} - m_f) q_f - \frac{1}{4} G^a_{\mu\nu} G^a_{\mu\nu}$$

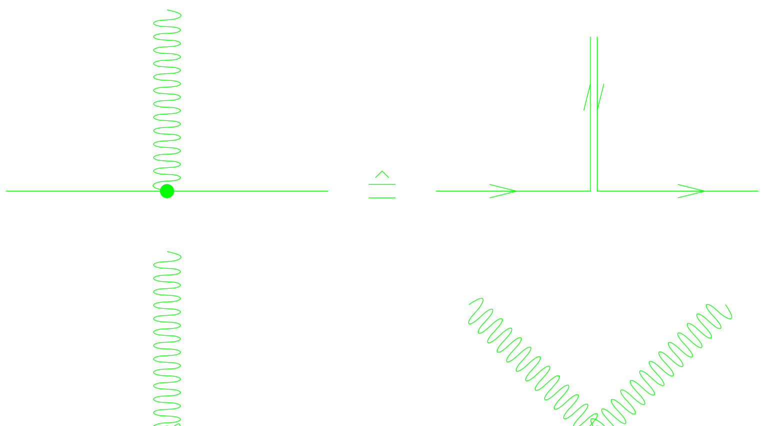
$$G^a_{\mu\nu} = \partial_\mu A^a_\nu - \partial_\nu A^a_\mu + g f^{abc} A^b_\mu A^c_\nu$$

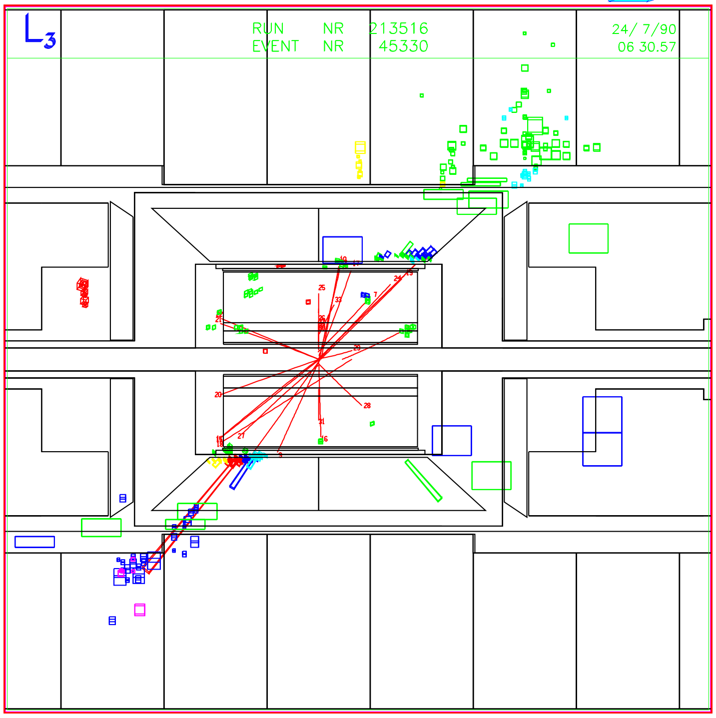
$$i\not{D} q = \gamma^\mu \left(i\partial_\mu + g A^a_\mu \frac{\lambda^a}{2} \right) q$$

- quantum mechanics: path integral (partition function)

$$Z = \int DA_\mu \ \ \mathrm{Det}(i\not{D} - m) \ \ \exp(i \int d^4x \mathcal{L}_g)$$

- dynamics in diagrams (Feynman)





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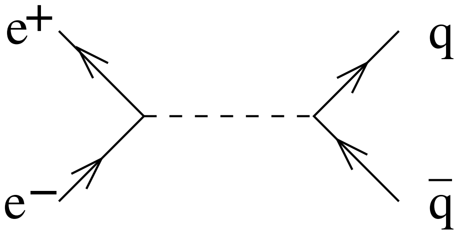
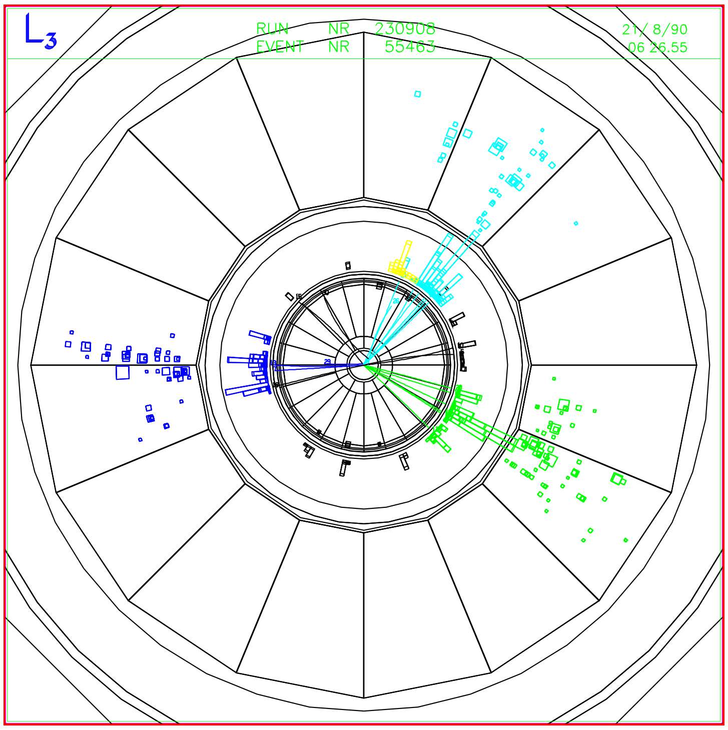
EVENT NR 45330
24/ 7/90
06 30.57



e+
e−
q
q̄



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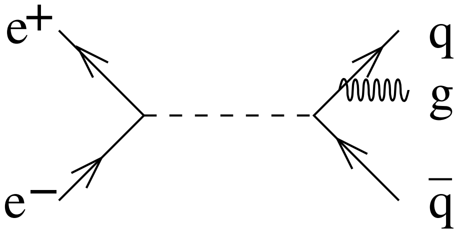
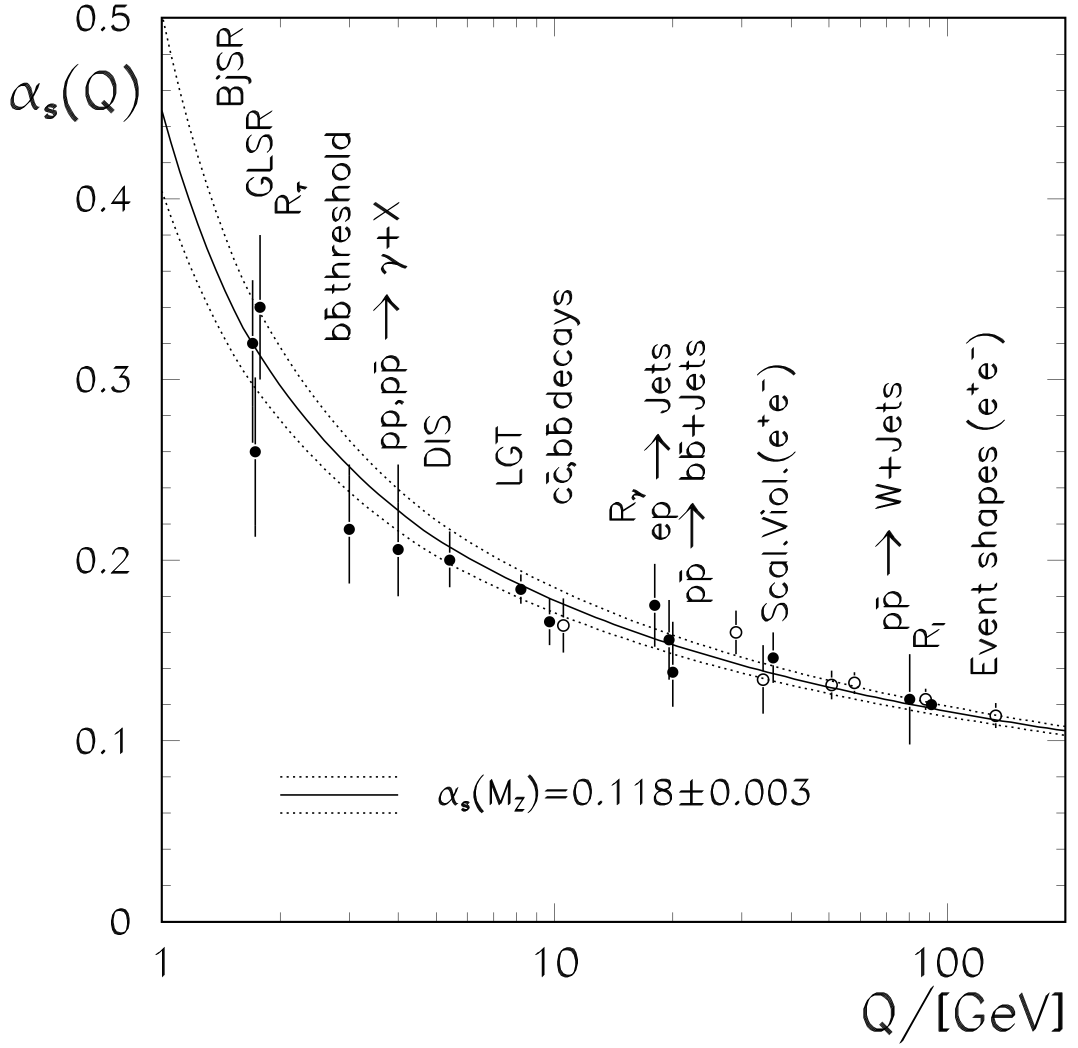
06 26.55



e+
e−
q
g
q̄



$\alpha_s(Q)$
BjSR
GLSR
R_τ
$b\bar{b}$ threshold
$pp,p\bar{p} \rightarrow \gamma+X$
DIS
LGT
$c\bar{c},b\bar{b}$ decays
R_γ
$ep \rightarrow$ Jets
$p\bar{p} \rightarrow b\bar{b}$+Jets
Scal.Viol.(e^+e^-)
$p\bar{p} \rightarrow W$+Jets
R_l
Event shapes (e^+e^-)
$\alpha_s(M_Z)=0.118\pm0.003$
0.5
0.4
0.3
0.2
0.1
0
1
10
100
Q/[GeV]

About units

- QCD is essentially parameter free: The lagrangian has a coupling constant, g, but no scale (if $m_q \simeq 0$).
- After renormalization g becomes scale dependent (g is traded for a scale parameter Λ)
- g is not a parameter, but QCD has an "urkilogramm"

$$\Lambda_{QCD} \simeq 200\,\mathrm{MeV} \simeq 1\,\mathrm{fm}^{-1}$$

  $\Lambda^{-1} \simeq 1\,\mathrm{fm}$ is the distance at which the coupling becomes strong
- In practice, perturbation theory breaks down at much shorter distances

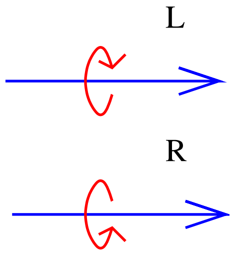
$$x \sim \Lambda_\chi^{-1} \sim 1\,\mathrm{GeV}^{-1} \sim 0.2\,\mathrm{fm}$$

- Note: $m_\rho \sim m_p \sim 1\,\mathrm{GeV}$, $r_p \sim 1\,\mathrm{fm}$.

Chiral Symmetry

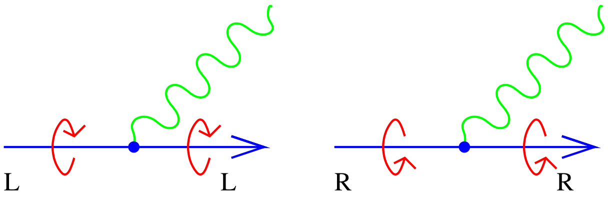
- define L/R fields

$$\psi_{L,R} = \frac{1}{2}(1 \pm \gamma_5)\psi$$

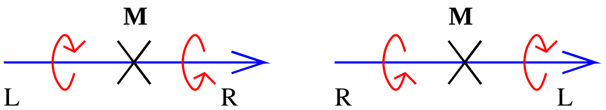


- fermionic lagrangian

$$\mathcal{L} = \bar{\psi}_L (i \not{D}) \psi_L + \bar{\psi}_R (i \not{D}) \psi_R$$



$$+ \bar{\psi}_L M \psi_R + \bar{\psi}_R M \psi_L$$



$$M = \text{diag}(m_u, m_d, m_s)$$

- $m_u = m_d = m_s$: Flavor symmetry $V \in SU(3)$

$$\psi_L \to V\psi_L, \qquad \psi_R \to V\psi_R$$

- $m_u = m_d = m_s = 0$: Chiral symmetry $(L, R) \in SU(3)_L \times SU(3)_R$

$$\psi_L \to L\psi_L, \qquad \psi_R \to R\psi_R$$

The Fate of Chiral Symmetry

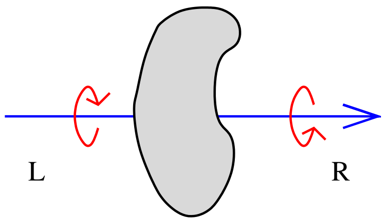
- Chiral symmetry implies parity doubling, e.g.

$$\begin{array}{rcc} & \rho(LL+RR) \quad \leftrightarrow & a_1(LL-RR) \\ \text{expt} & m_\rho = 770\,\text{MeV} & m_{a_1} = 1230\,\text{MeV} \end{array}$$

too large in order to be explained by quark masses $m_u, m_d \sim 10\,\text{MeV}$

- chiral symmetry is spontaneously broken

$$\langle \bar\psi_L \psi_R + \bar\psi_L \psi_R \rangle \simeq -(230\,\text{MeV})^3$$

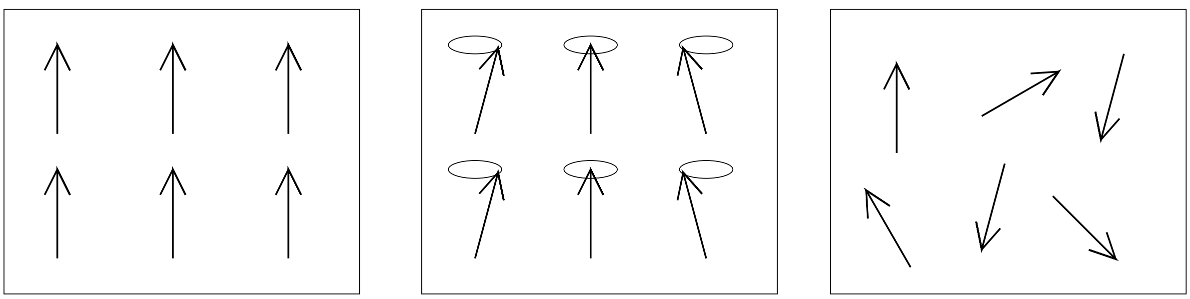


- consequences: dynamical mass generation $m_Q = 300\,\text{MeV} \gg m_q$

$$\begin{array}{cccc} m_p = & 890\,\text{MeV} & + & 45\,\text{MeV} \\ & (\text{QCD},\ 95\%) & & (\text{Higgs},\ 5\%) \end{array}$$

- existence of Goldstone modes

$$\langle \bar\psi^f_L \psi^g_R \rangle = B\,(\sigma + i\vec\tau\vec\pi)^{fg}$$

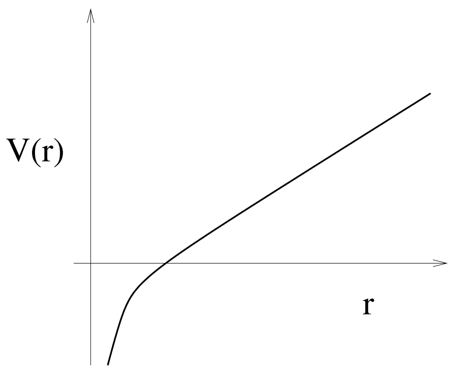


ordered | spin wave | disordered

chiral symmetry broken | pion | chiral symmetry restored

Confinement

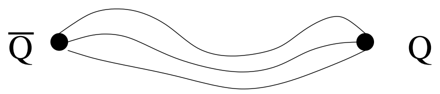
- heavy quark potential



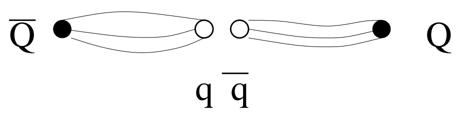
$$V(r) \sim \frac{\alpha}{r} + kr$$

$$k \sim 1\mathrm{GeV/fm}$$

- picture: gluon field channeled into flux tubes

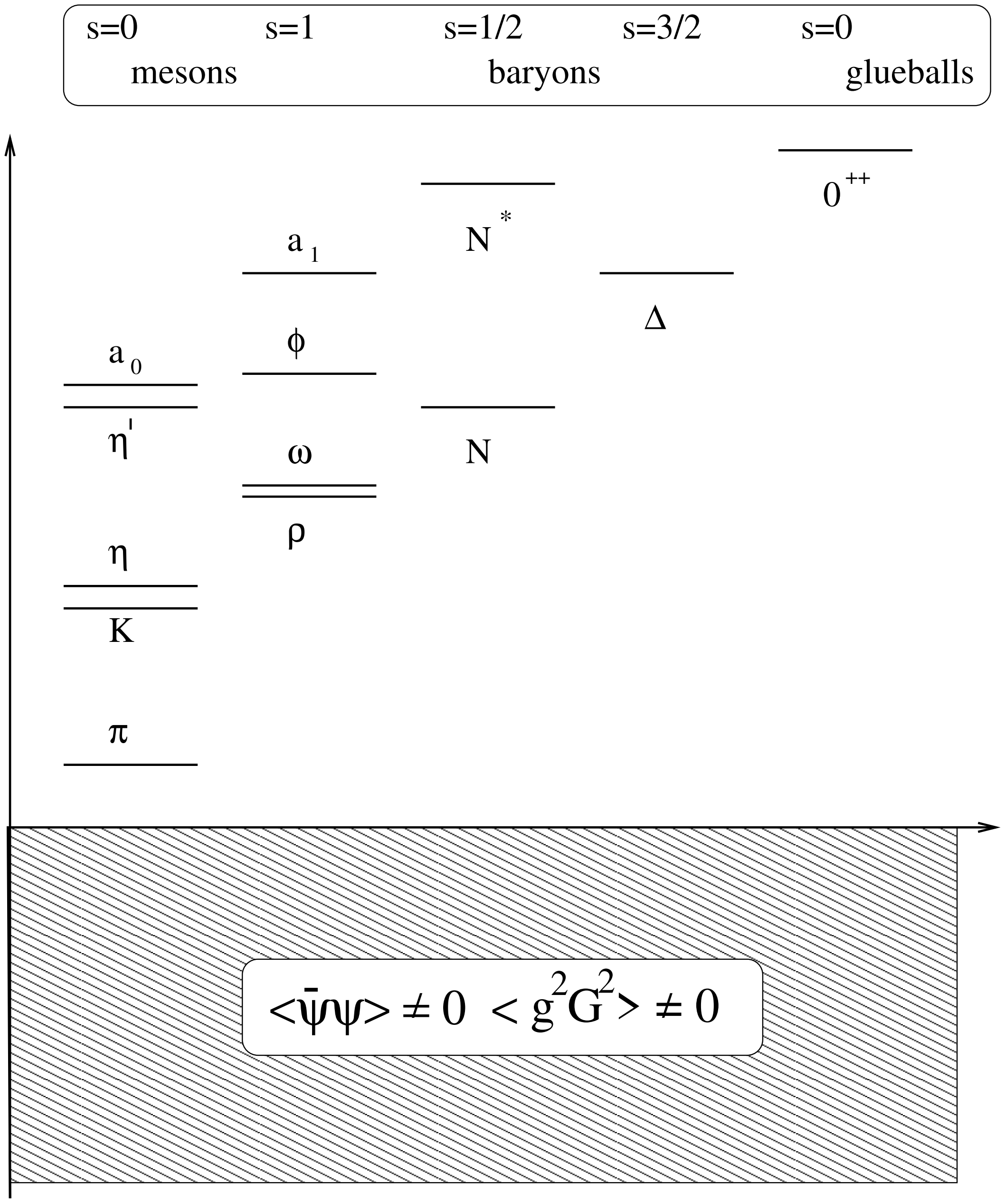


- light quarks: string can break



$$J/\psi(\bar{c}c) \to D(\bar{c}u) + \bar{D}(\bar{u}c)$$

QCD Ground State



Chiral Symmetry restoration

- vacuum expectation value

$$\langle \sigma \rangle = f_\pi \qquad f_\pi = 93\,\mathrm{MeV}$$

- thermal pion fluctuations

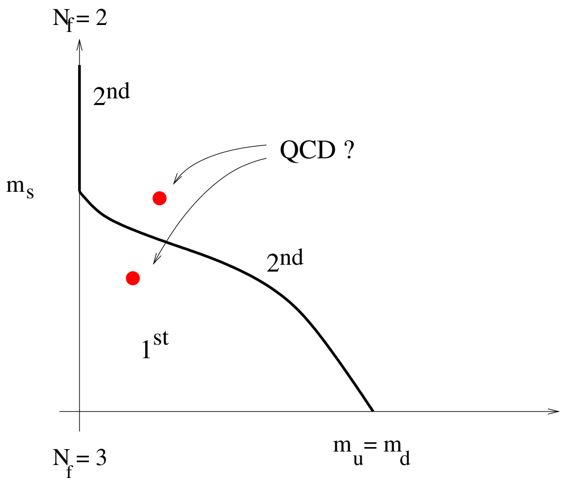
$$\langle \sigma \rangle = \langle \sigma \rangle|_{T=0} \left\{ 1 - \frac{N_f^2 - 1}{3N_f} \left(\frac{T^2}{4f_\pi^2} \right) + \ldots \right\}$$

$$\to T_c \sim \sqrt{8} f_\pi \sim 200\,\mathrm{MeV}$$

- universality: critical behavior near 2^{nd} order phase transition is independent of microscopic details

$$N_f = 2 \;\; 2^{\mathrm{nd}}\,\mathrm{order},\; O(4)\,\mathrm{critical\ exponents}$$

$$N_f = 3 \;\; 1^{\mathrm{st}}\,\mathrm{order}$$



The QCD phase transition